

# Experimentally Driving the Inverse Energy Cascade in a Dipole Confined Plasma

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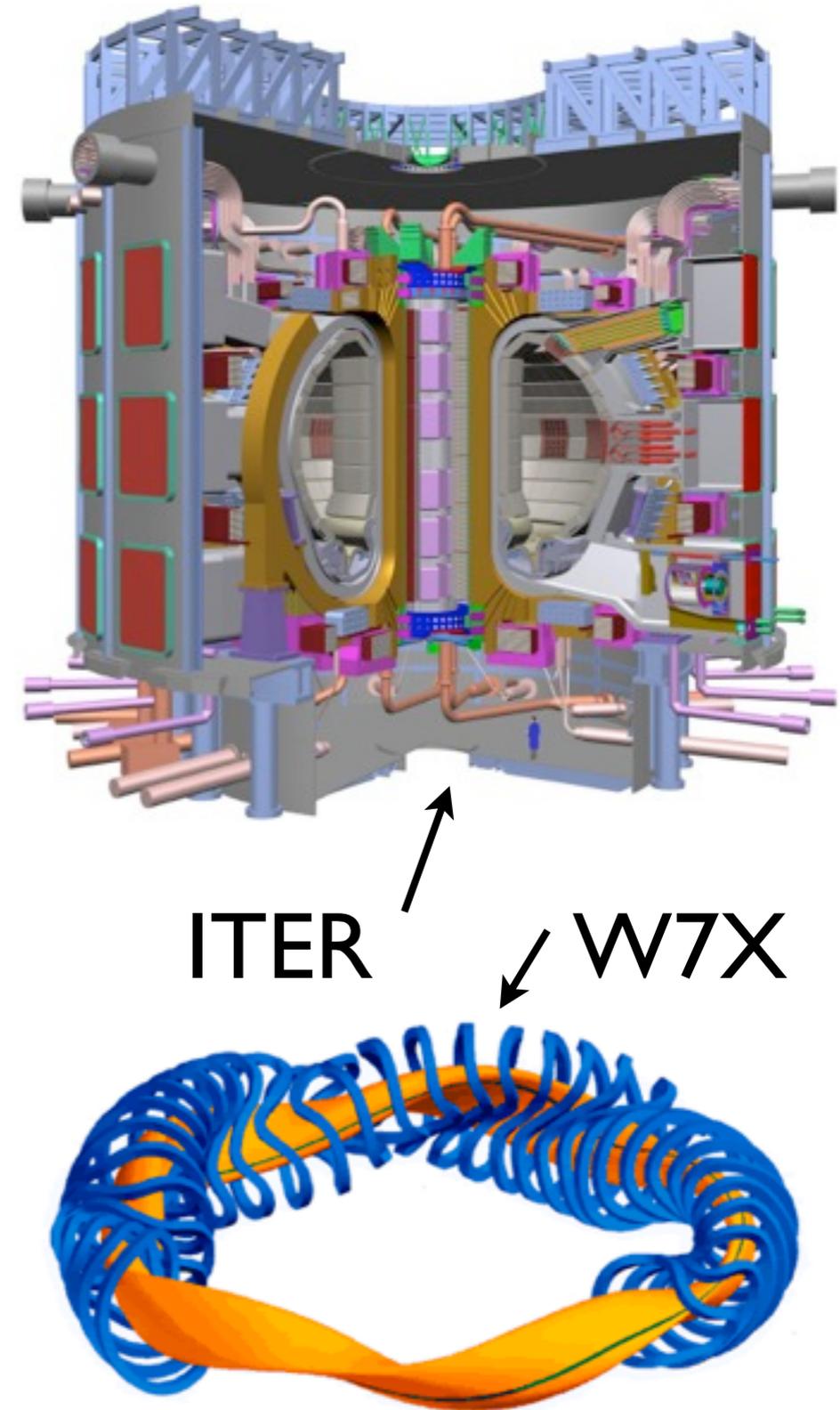
# Outline

- Dipoles within the topological zoo
- The Collisionless Terrella Experiment (CTX)
- Introduction of turbulence on CTX
- Turbulence, what is it? What does this mean?
- Modifying the turbulent spectrum



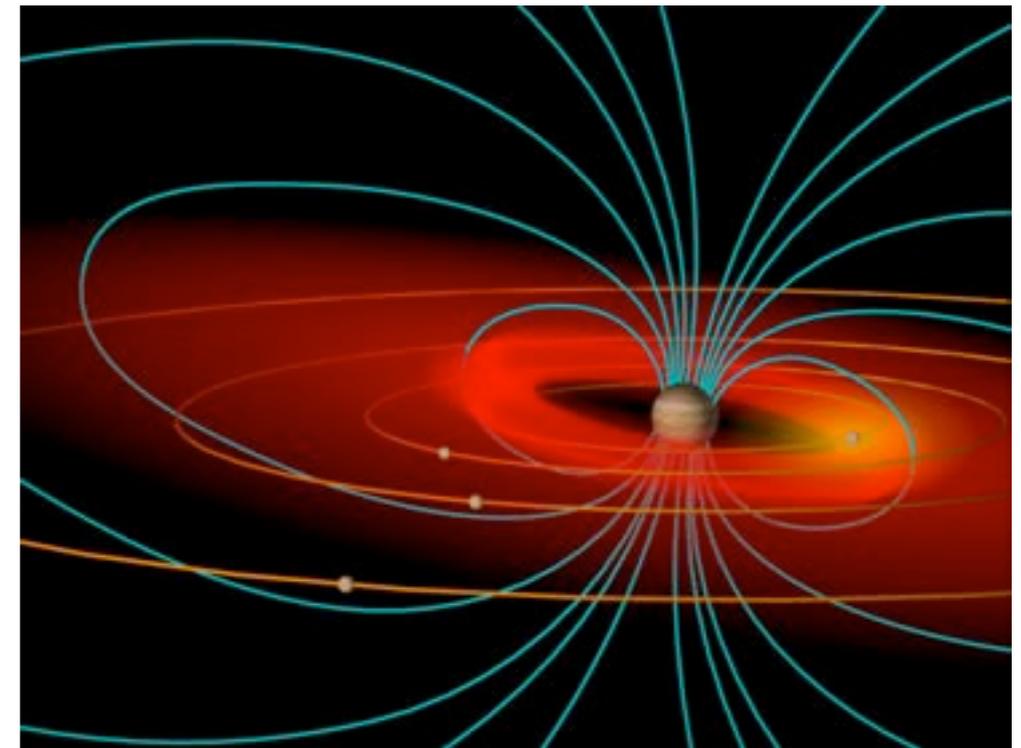
# Magnetic Fusion Topology

- Most common magnetic geometries for plasma confinement employ magnetic surfaces
- the Tokamak is the most common and relies on the plasma current to create magnetic surfaces
- the Stellarator uses shaped magnetic coils to create magnetic surfaces, plasma current not necessary
- other options like the Reverse Field Pinch (RFP) or Field Reverse Configuration (FRC) also exist

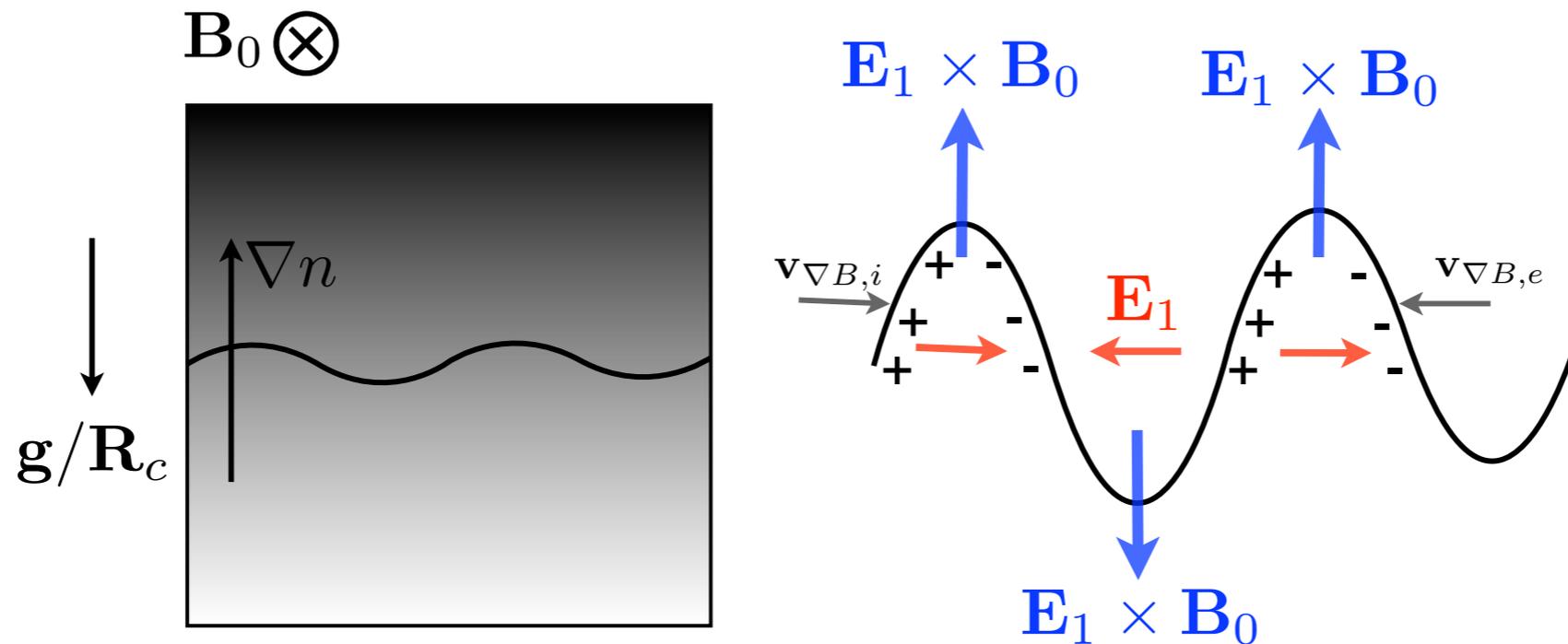


# The Dipole Option

- Nature's choice for magnetic confinement.
- The field created by a magnetic dipole is a purely poloidal. All field lines close on themselves, creating "flux tubes" or "flux ropes" rather than magnetic surfaces.
- This experimental geometry has several advantages including: great diagnostic access, simple coil design, the possibility of advanced fuels, and bi-dimensional (2D) physics

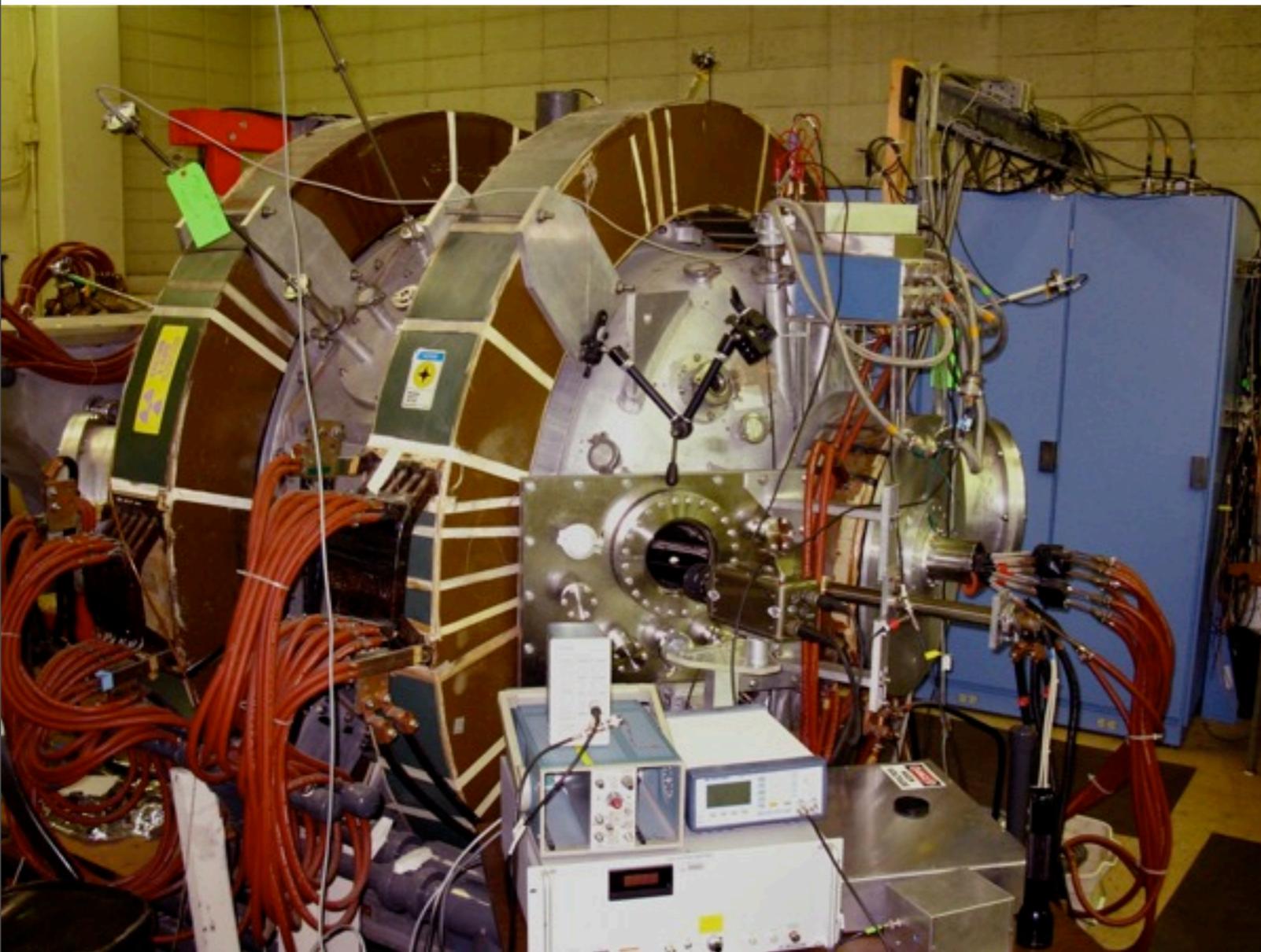


# Interchange Instability is the Primary Dipole Instability



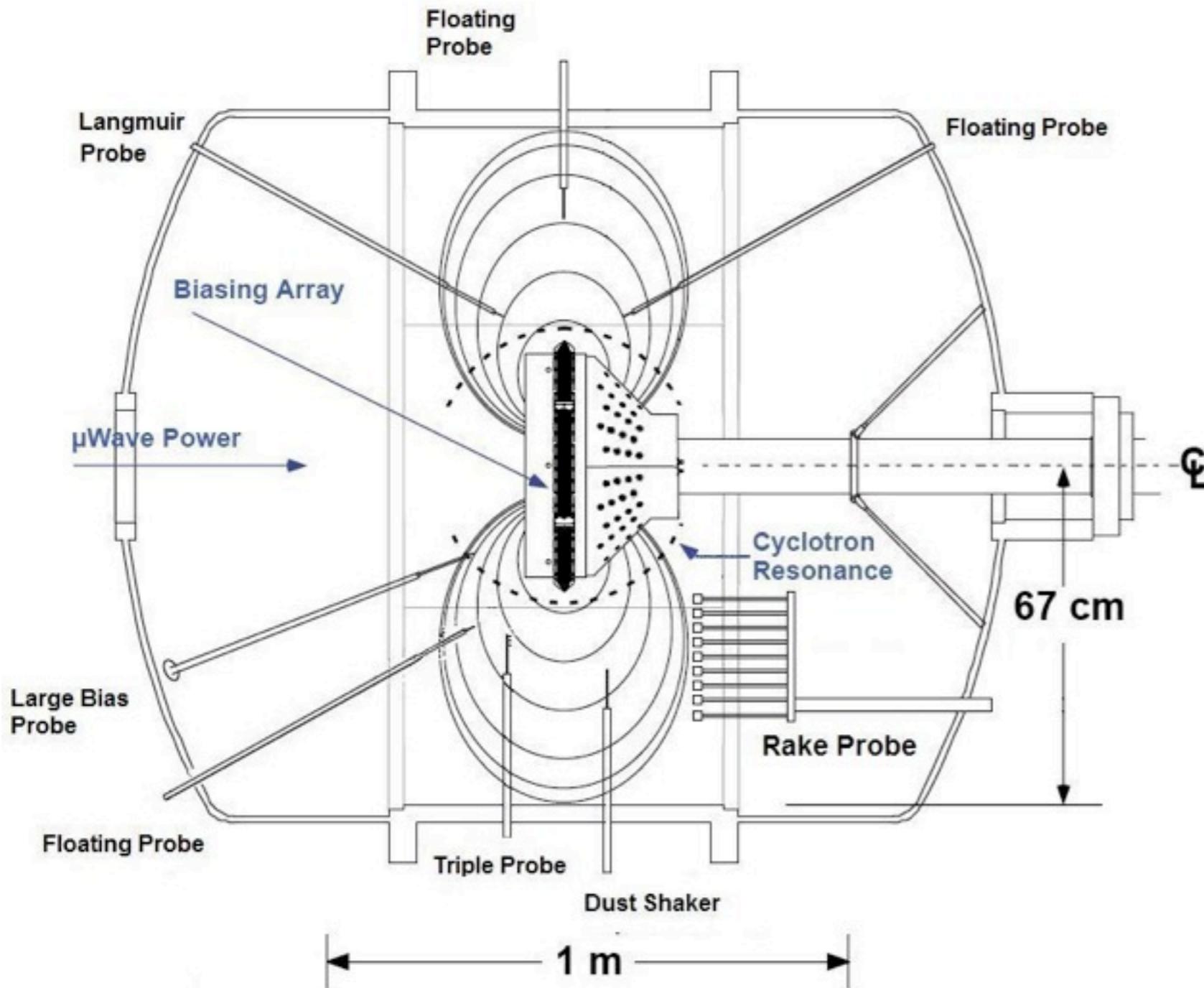
- Analogous to the gravitational Rayleigh-Taylor instability
- Similar to turning a cup of water upside down
- No field-aligned dynamics

# Creating Plasma on the Collisionless Terella Experiment



- Hydrogen plasma created with Electron Cyclotron Resonance Heating. (ECRH)
- 1.0kW of RF waves with  $f=2.45\text{GHz}$  injected.
- $B_0=875\text{G}$ , resonance at  $L=27\text{cm}$  creates ring of deeply trapped, hot electrons.
- Base Vacuum Pressure  $\approx 1-2 \times 10^{-7}$  Torr

# Diagnostics to Characterize Plasma Turbulence and Flow



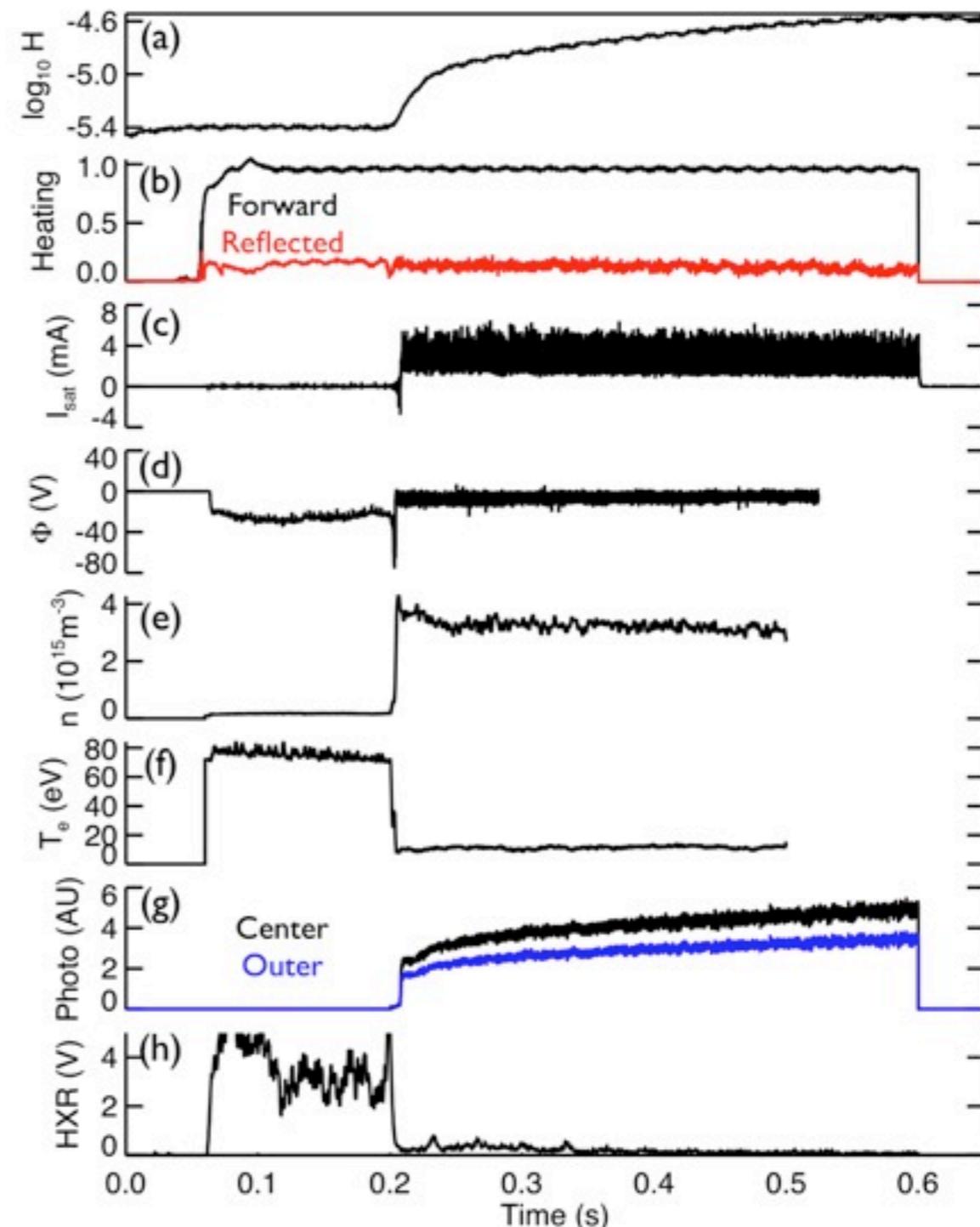
- 3 Langmuir Probes
- Triple Probe Array
- $I_{\text{sat}}$  Probe
- Bias Probe
- Equatorial Array
- 16 Point, 31 Tip Radial Transport Rake Probe Array
- Polar Imager (Array of 96 Gridded Energy Analyzers)



# Plasma Discharge Enters Turbulent Regime with Sufficient Fueling\*

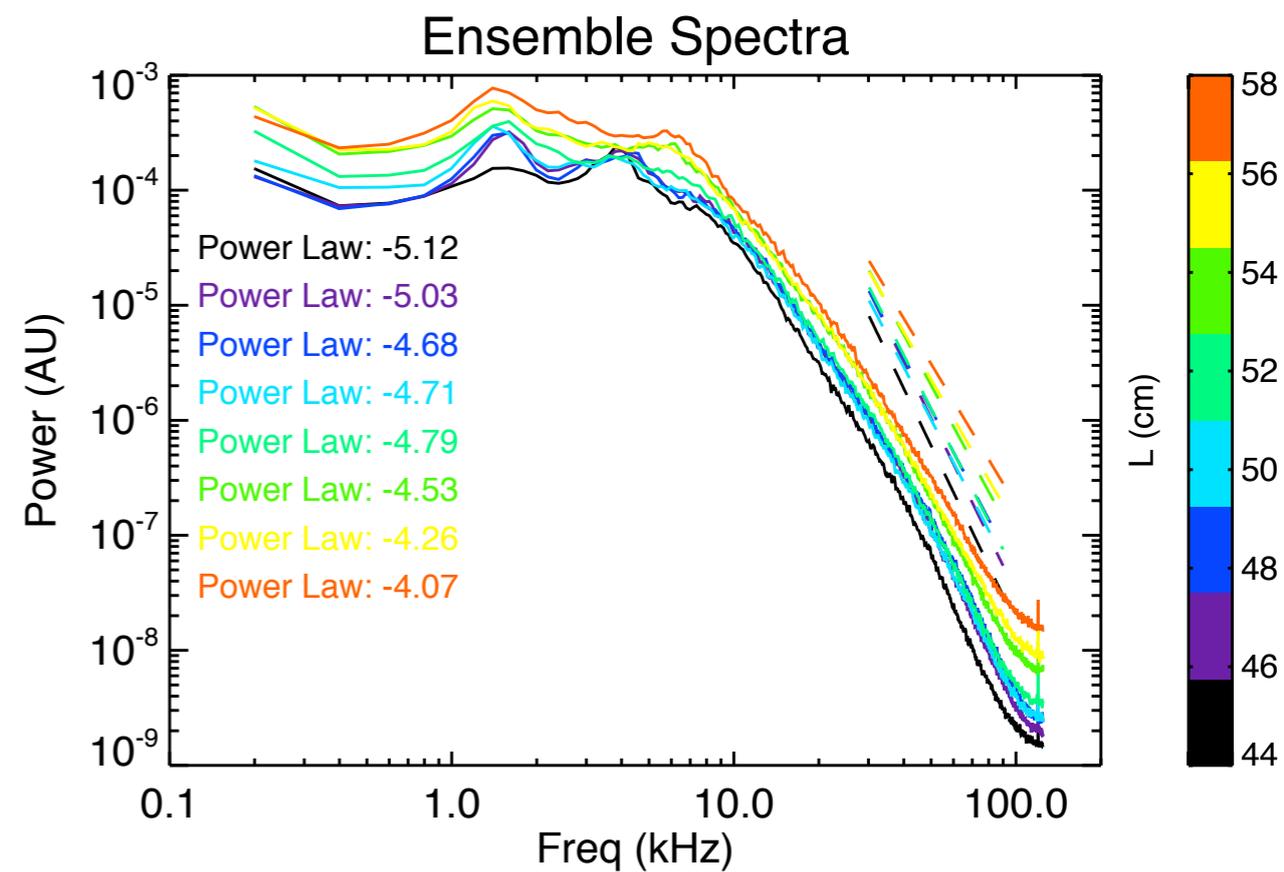
- Transition to high density turbulent regime at  $t = 0.2$  s.
- Temperature drops an order of magnitude.
- The edge density increases to  $n \sim 3-4 \times 10^{15} \text{ m}^{-3}$ .
- Visible light increases dramatically and hard X-ray production decreases.

\*N.A. Krall *Phys. Fluids* **9** (4)(1966)



# High Density Single Point Measurements Look Turbulent

- After transition floating potential fluctuations exhibit broad power law spectrum.
- Two peaks observed at roughly 2 kHz and smaller peak at 4 kHz.



# What Does Turbulence Mean?

- Turbulence implies highly stochastic, nonlinear, multi-scale dynamics
- Occurs in most flowing media: water, air, plasma
- A quote attributed to British physicist Horace Lamb is, “I am an old man now, and when I die and go to heaven there are two matters on which I hope for some enlightenment. One is quantum electrodynamics, and the the other is the turbulent motion of fluids. And about the former I am rather optimistic.”

# Enstrophy plays a large factor in differences between 3D and 2D

- Enstrophy is the mean square vorticity.
- Not a conserved quantity in 3D b/c of vortex stretching. (think of water draining through a funnel)
- Conserved in 2D in the incompressible, inviscid limit.
- In the Fourier representation we can see the relationship between energy and enstrophy.

$$\omega = \nabla \times \vec{v}$$

$$\epsilon = \langle \omega^2 \rangle$$

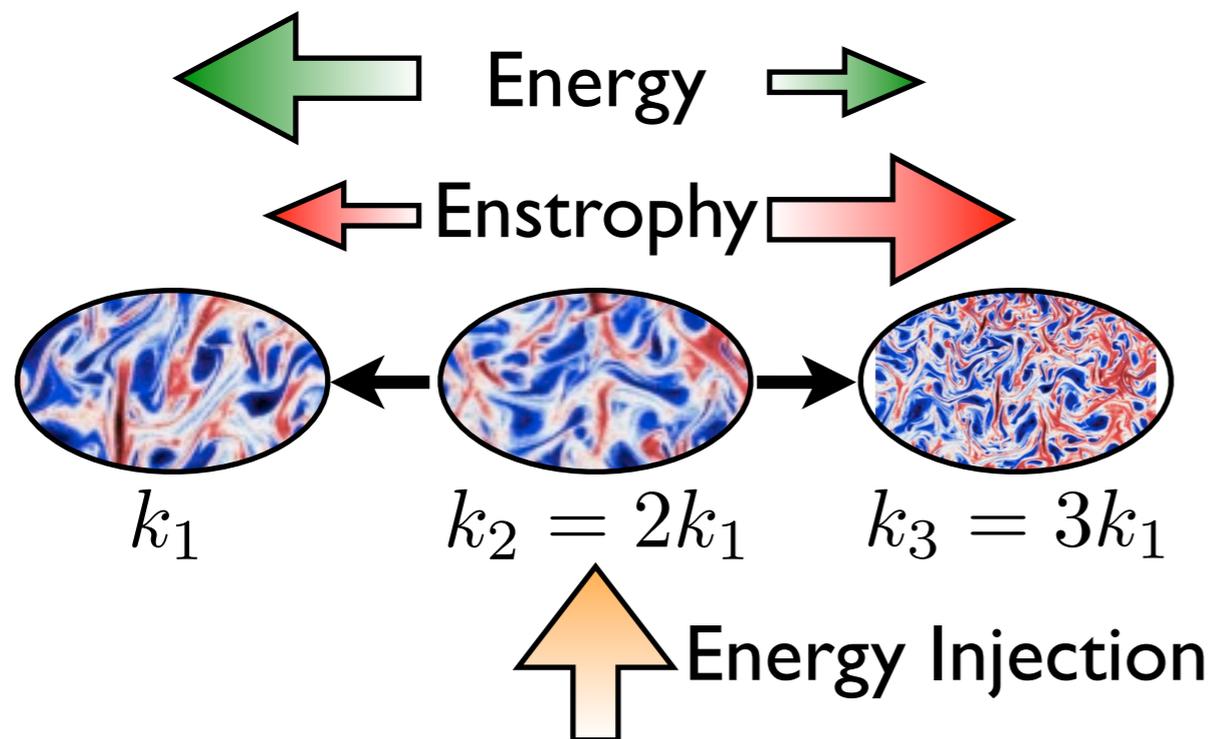
$$E = \int_0^\infty S(k) dk$$

$$\epsilon = \int_0^\infty k^2 S(k) dk$$

$$\epsilon = k^2 E$$

# Implications of Enstrophy Conservation

- Let's look at the effects this has with a simple example:



$$\vec{k}_1 + \vec{k}_2 = \vec{k}_3, |k_1| < |k_2| < |k_3|$$

$$\delta E_1 + \delta E_2 + \delta E_3 = 0$$

$$\delta \epsilon_1 + \delta \epsilon_2 + \delta \epsilon_3 = 0$$

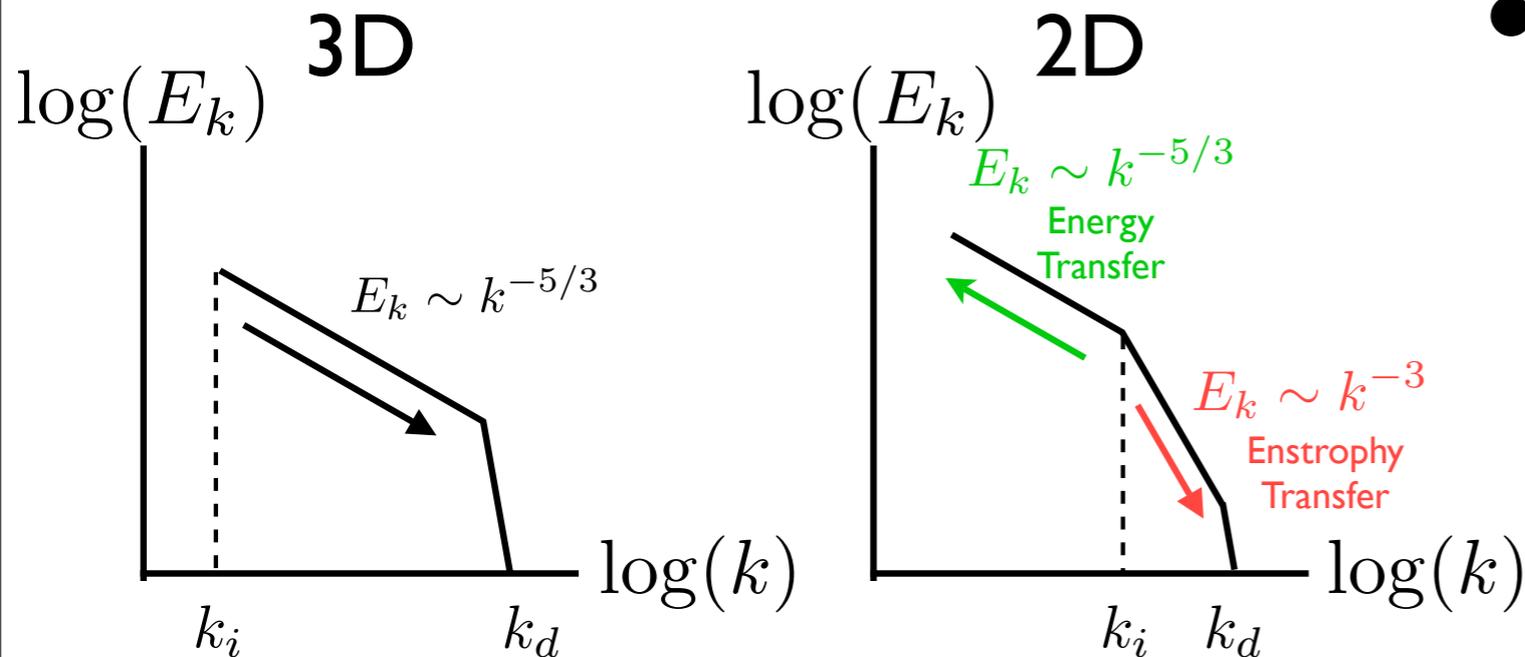
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$$k_1^2 \delta E_1 + k_2^2 \delta E_2 + k_3^2 \delta E_3 = 0$$

Do the algebra...

$$\frac{\delta E_1}{\delta E_3} = \frac{5}{3} \quad \frac{\delta \epsilon_1}{\delta \epsilon_3} = \frac{5}{27}$$

# This Creates a Dual Cascade

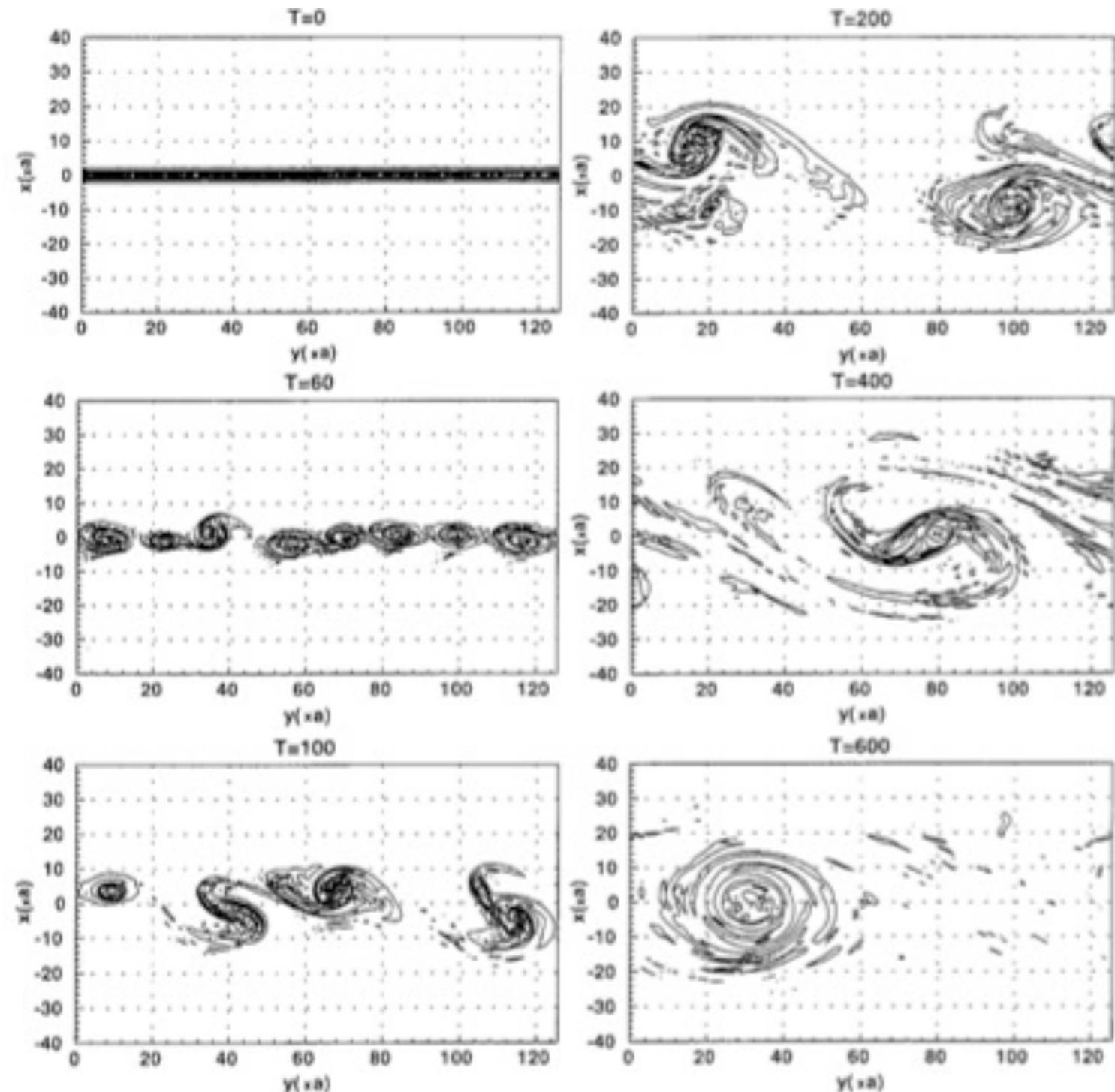
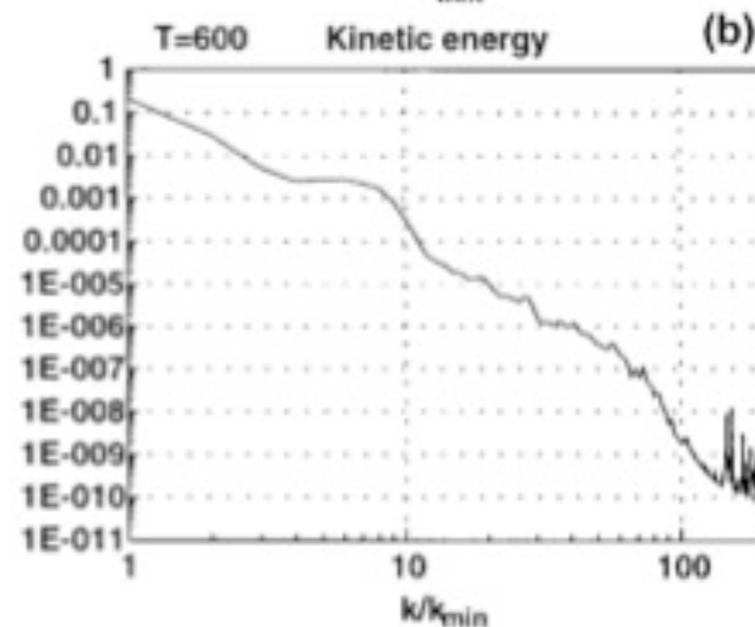
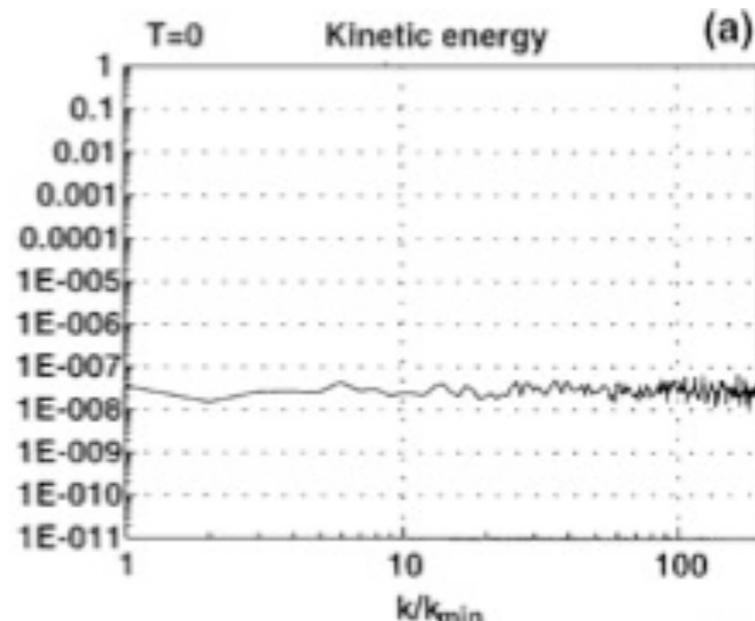


- So we see as a result of reducing the dimensionality, we have another conserved quantity (enstrophy), which in turn creates dual cascades.

- Enstrophy exhibits a forward cascade from large to small spatial scales. (small to large k-values)  $k^{-3}$
- Energy has a backward cascade, the so-called inverse energy cascade, from small spatial scales to larger ones. (large to small k-values)  $k^{-5/3}$

# An Inverse Cascade Example in a 2-D Kelvin-Helmholtz Simulation

\*A. Miura, PRL **83**, 8 (1999)



# Are We Observing the Inverse Energy Cascade?

- We have reported that in the high density turbulent regime the observed fluctuations can be recreated through spatially broad and temporally chaotic signals\*
- This is consistent with an inverse energy cascade (pooling at large spatial scales), but we aim to inject energy at finer scales and watch the spectrum evolve
- The goal is to experimentally verify the inverse energy cascade through application of an electrostatic perturbation

\*Grierson, et al. Phys. Plasmas **16**, 5 (2009)

# How can we most effectively couple to the plasma?

- Using a bias probe the data indicate the perturbation does modify the plasma. However, limited by a single probe and the amount of power that can be applied
- Increasing the probe area would allow more power to be applied
- Desire ability to control the shape of the perturbation through broader spatial coverage

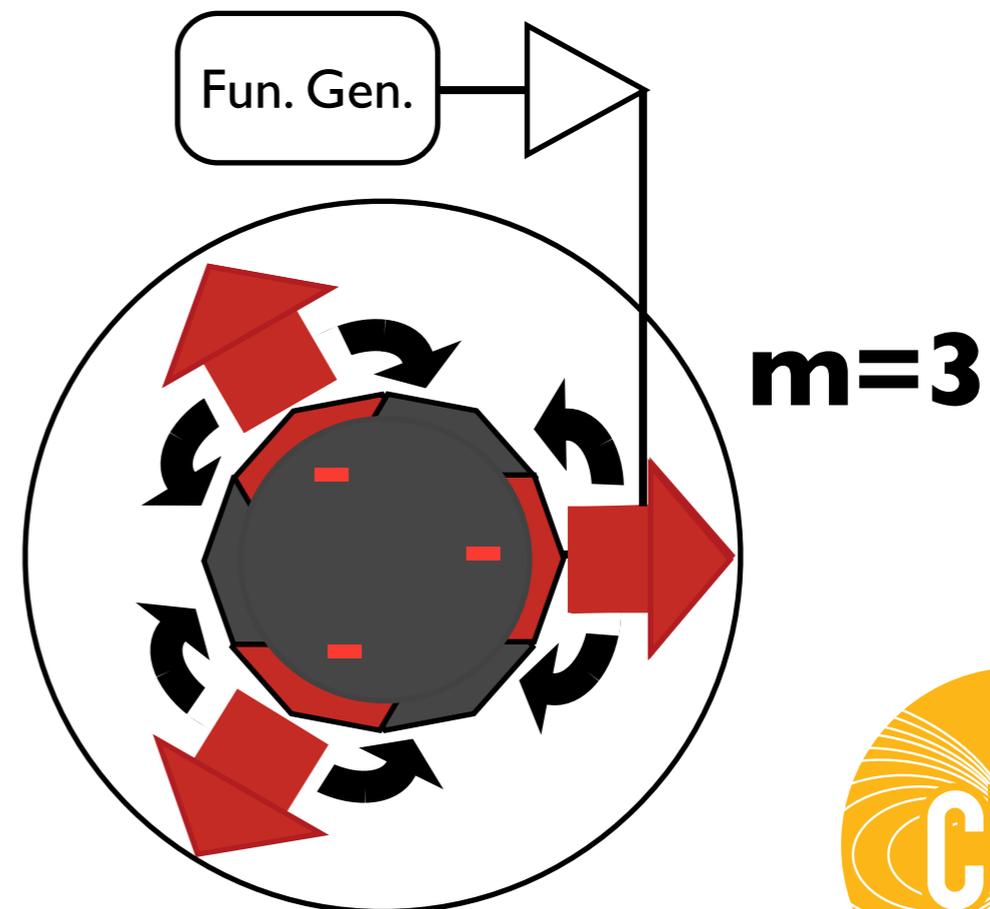
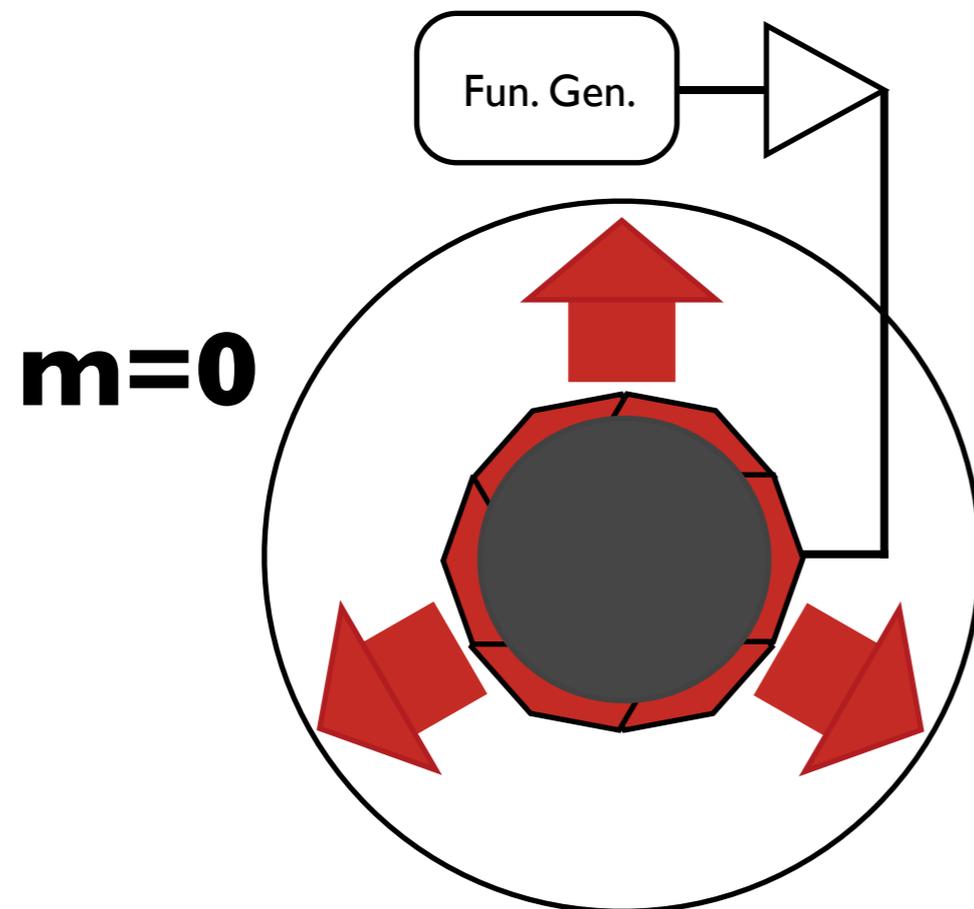
# Equatorial Biasing Array

- Upgrade existing array and begin experiments
- Increase number of segments to twelve
- See symmetry breaking effect,  $m=0$  has  $\sim 3x$  the current with  $2x$  the number of meshes biased

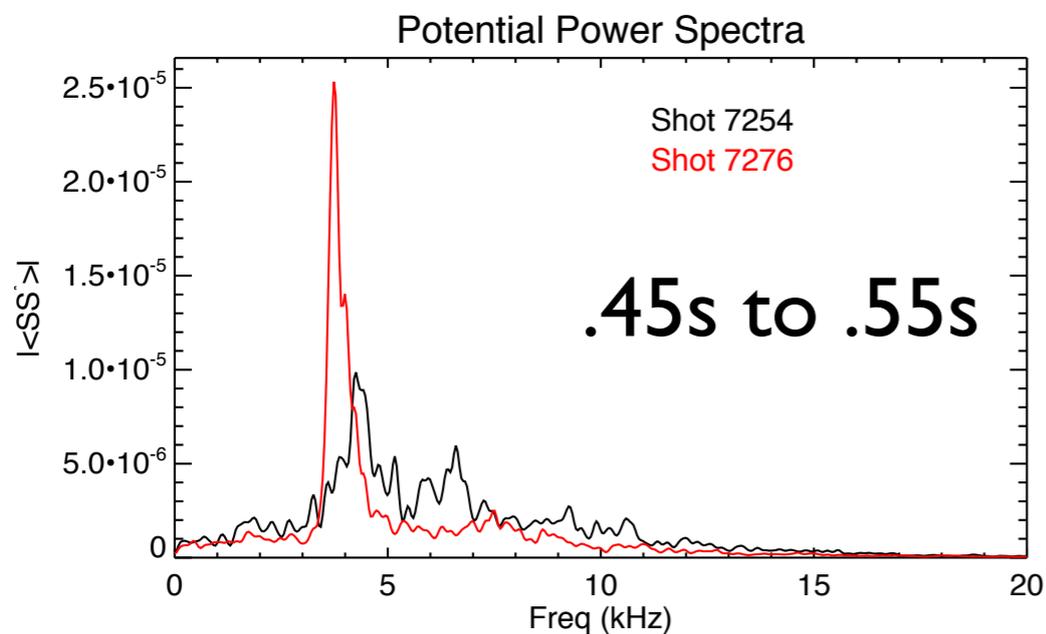
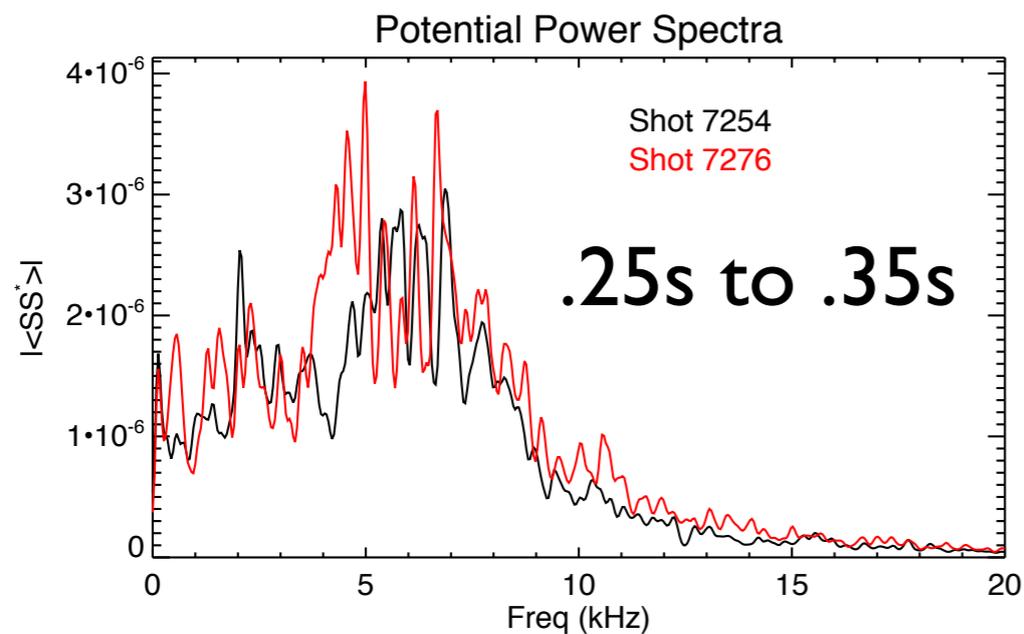


# Driving Azimuthal vs. Radial Currents through Varied Mode Numbers

- When axisymmetry is broken currents can flow azimuthally in addition to radially
- Equatorial bias array designed to test mode numbers=0,1,2,3,6

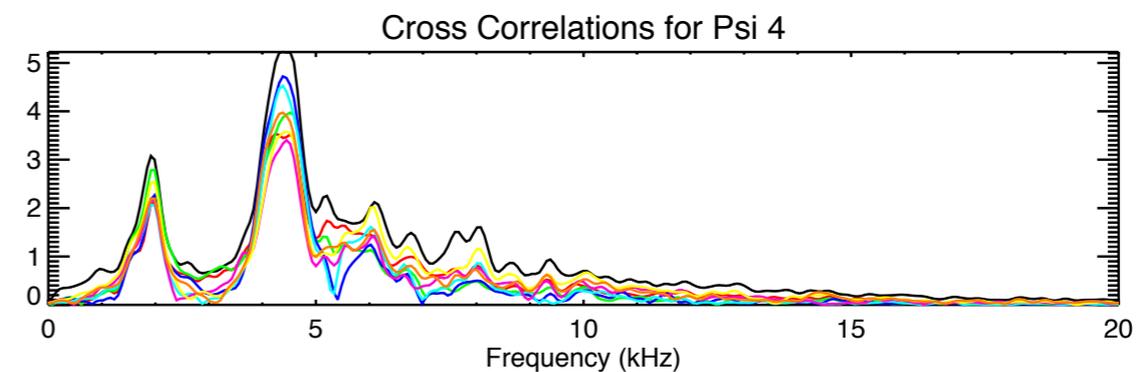
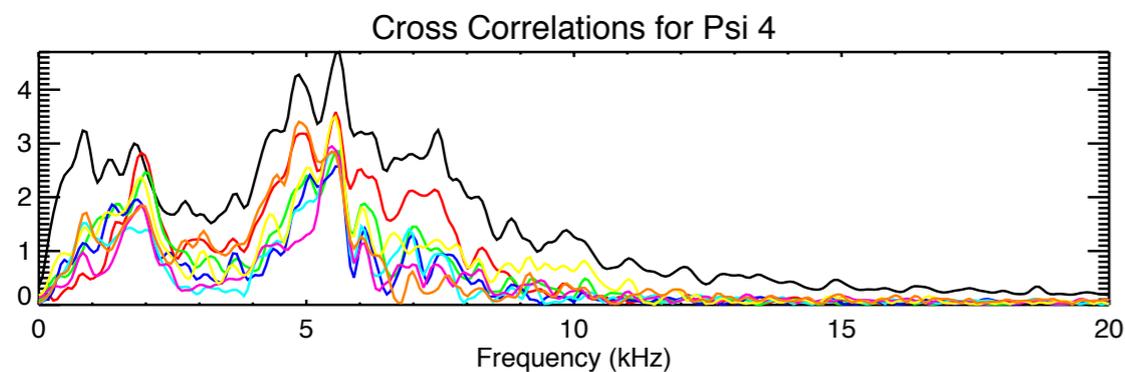
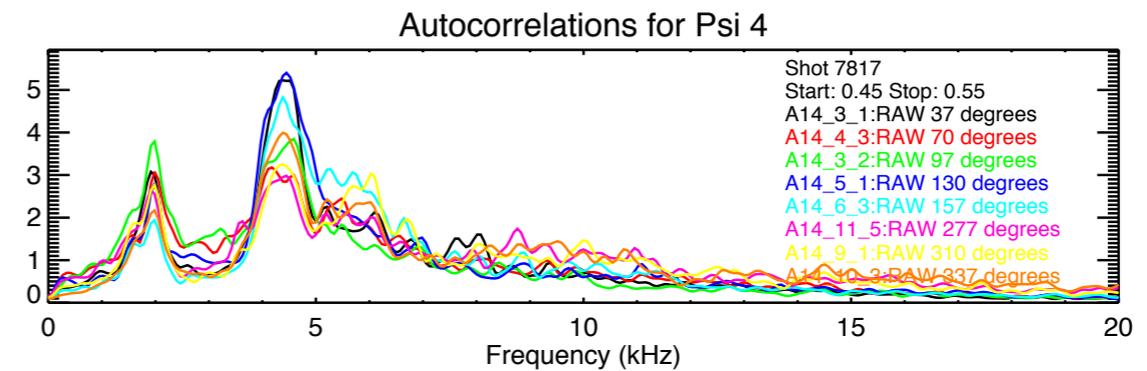
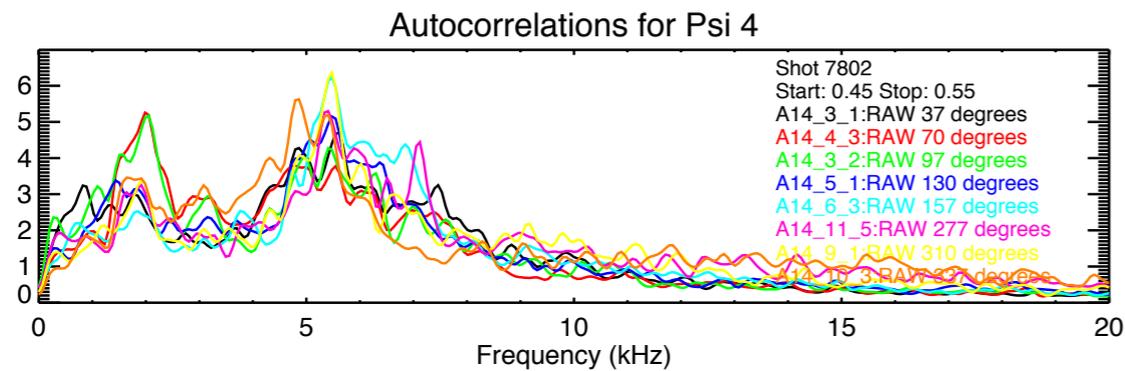


# Equatorial Bias Excites Quasi-Coherent Mode



- Shots similar before bias, but see the quasi-coherent mode when bias is applied
- 7254: High Density
- 7276: -500V  $m=6$  bias triggered from .35 to .55s, full bias around .45s

# Mode Seen in Polar Imager Correlations

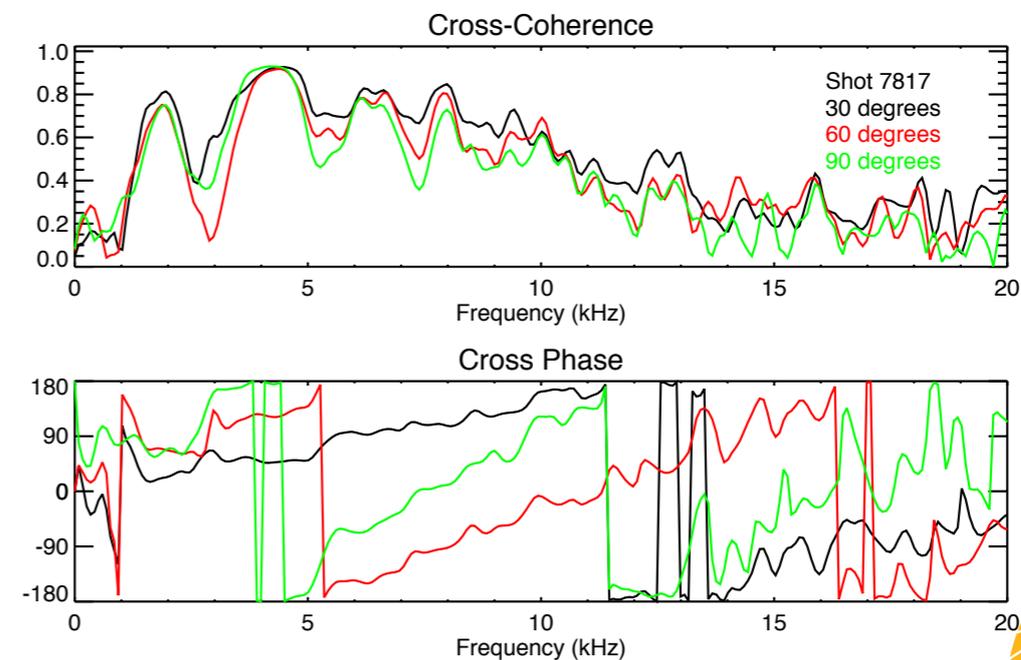
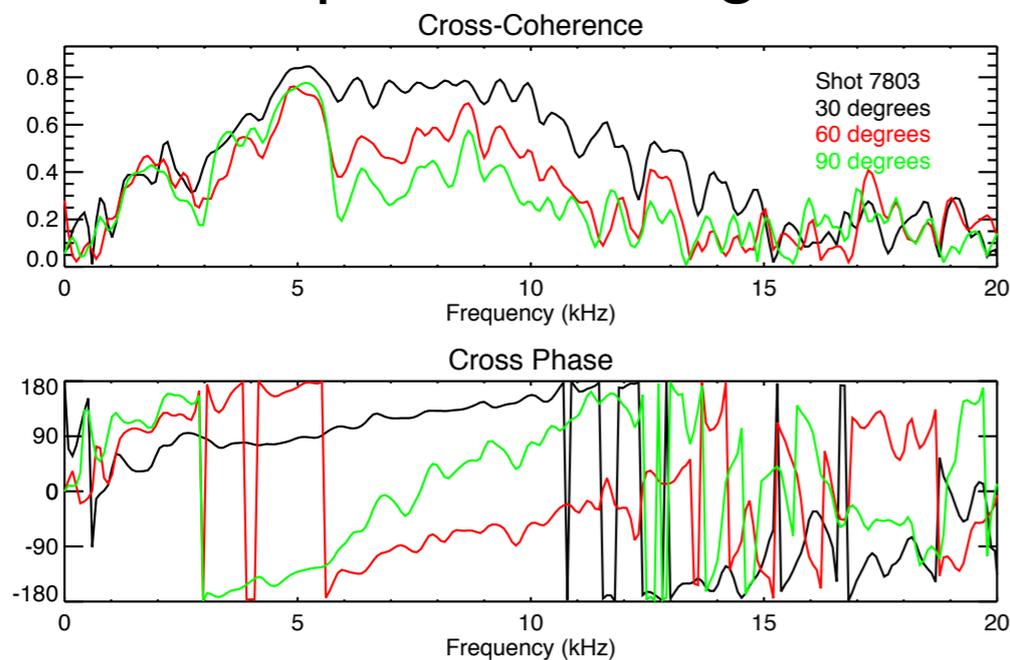


**Unbiased**

**m=6 Bias -400V**

# Bias Increases Coherence in Potential Probes

- Unbiased shot displays decreased coherence length beyond dominant mode.
- Multiple modes present in shot with nonaxisymmetric bias.
- Increased coherence length evident through broad range of frequencies during bias.



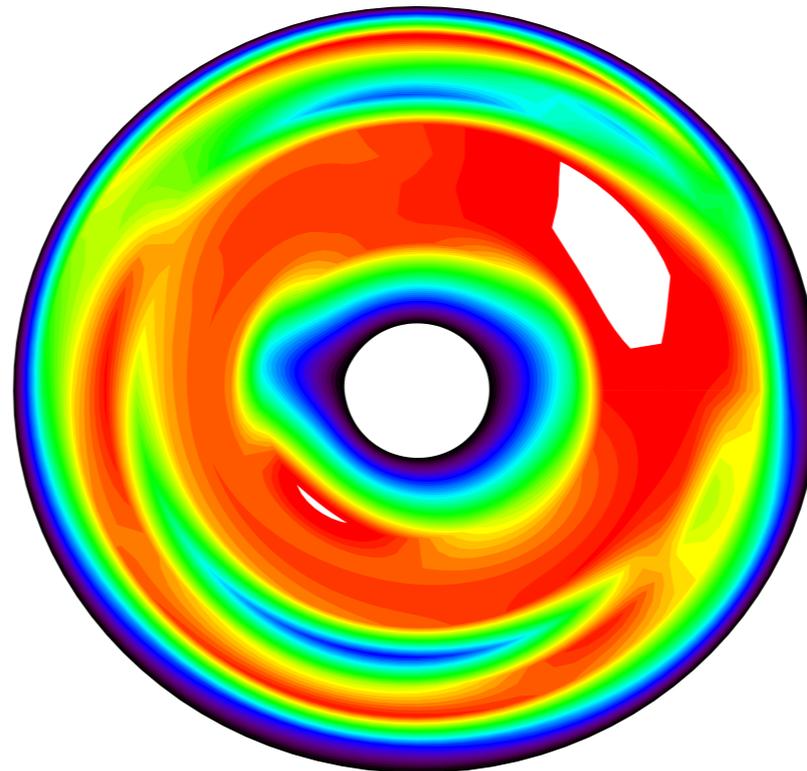
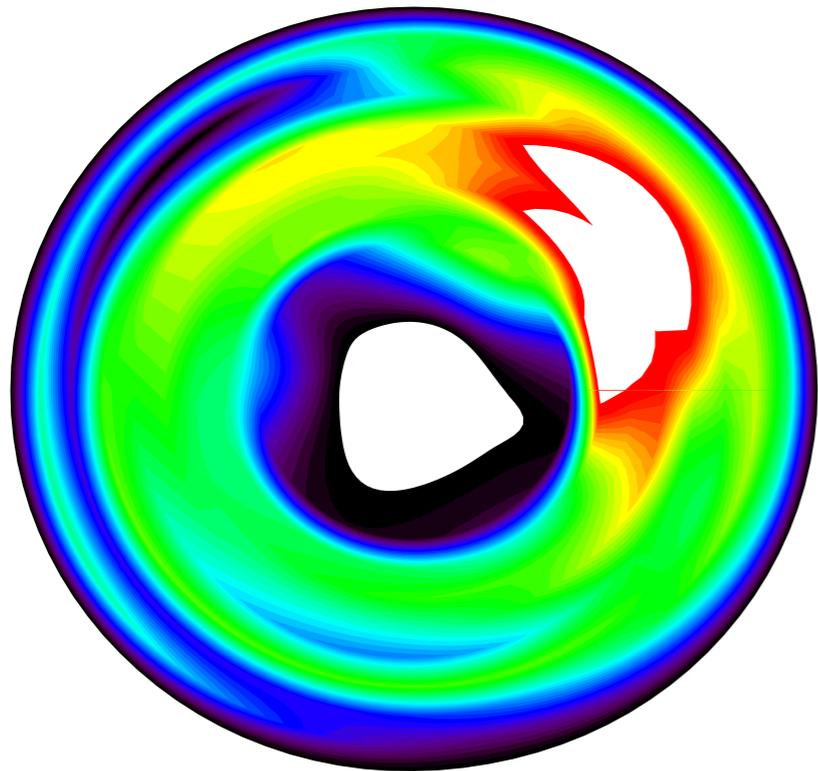
**Unbiased**

**m=6 Bias -400V**

# Coherence of Primary Mode Increases During Bias

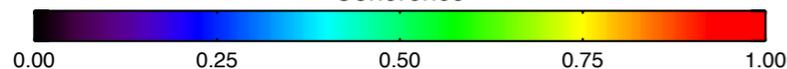
**Unbiased**

**m=3 Bias -400V**

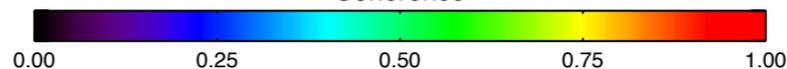


- Strong azimuthal increase in coherence
- Coherence length also broadens radially

Reference Detector: A14\_3\_1:RAW  
Shot 7282 Start: 0.500000 Stop: 0.550000  
Frequency Band=[4.30000,4.90000]  
Coherence

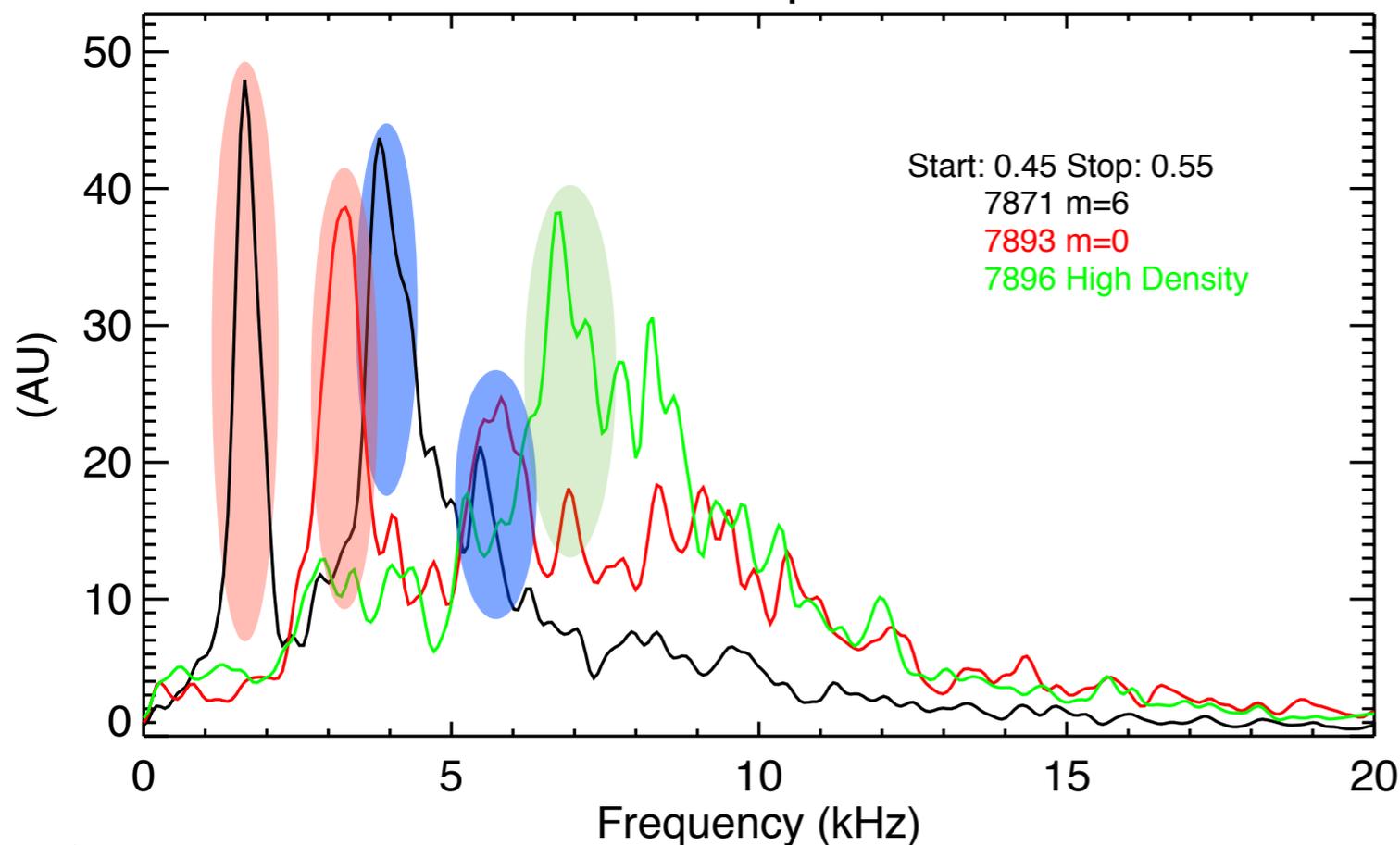


Reference Detector: A14\_3\_1:RAW  
Shot 7187 Start: 0.500000 Stop: 0.550000  
Frequency Band=[3.10000,3.70000]  
Coherence



# Nonaxisymmetric Bias Excites Inverse Energy Cascade

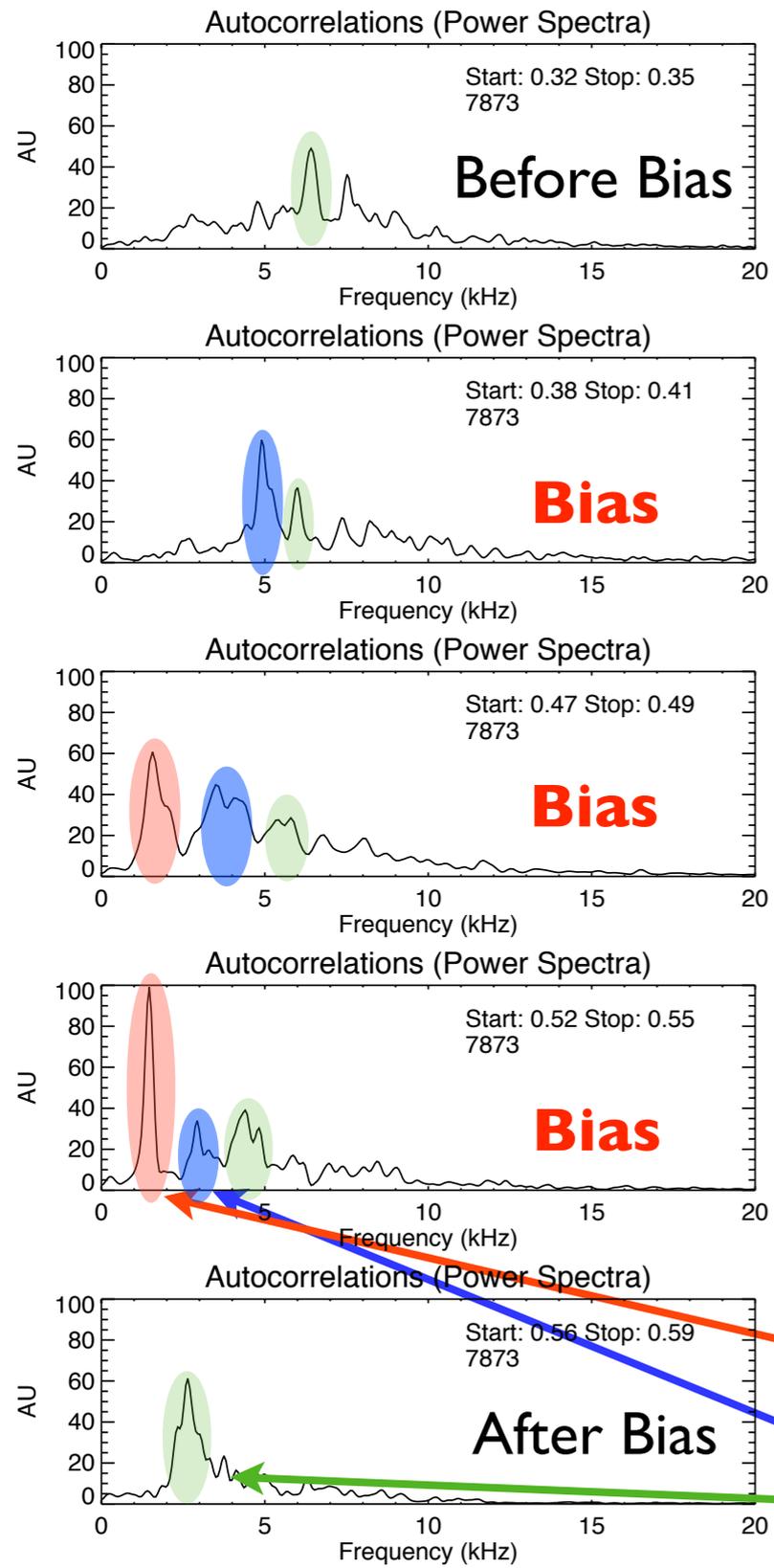
**m=1 modes** **m=2 modes** **m=3 mode**  
Power Spectra



- Floating potential spectra
- Structure of unbiased dominant mode m=3 (not very coherent)
- Axisymmetric mode shows frequency upshift
- Nonaxisymmetric bias displays clear excitation of lower mode numbers
- RMS unchanged by bias



# Potential Time Evolution During Bias



**Time**

**m=1 mode**  
**m=2 mode**  
**m=3 mode**

- We see frequency shift and mode amplitude jump after bias is triggered, and then decay once bias is removed
- Unbiased shots display a growing amplitude until saturates with slowing of principal mode (neutral drag)



# Spectral Analysis Can Reveal Nonlinear Coupling of Scales

- Standard FFTs and Spectrograms employed.
- Bicoherence also useful in quantifying three-wave coupling within a signal.

$$S(t) \xrightarrow{FFT} \hat{S}(\omega)$$

$$\langle A \rangle = \frac{1}{M} \sum_{i=1}^M A_i$$

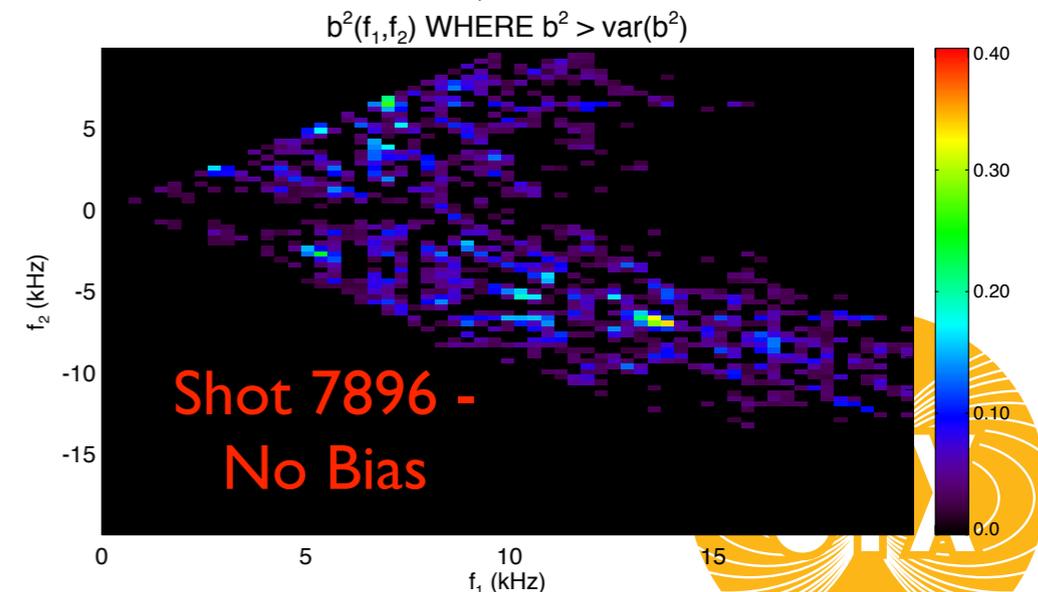
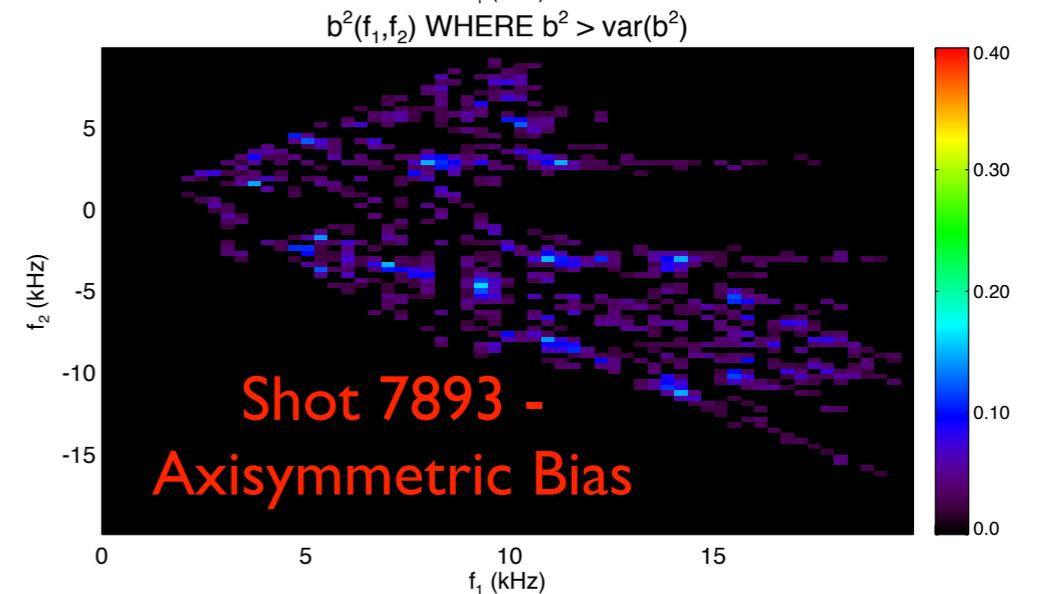
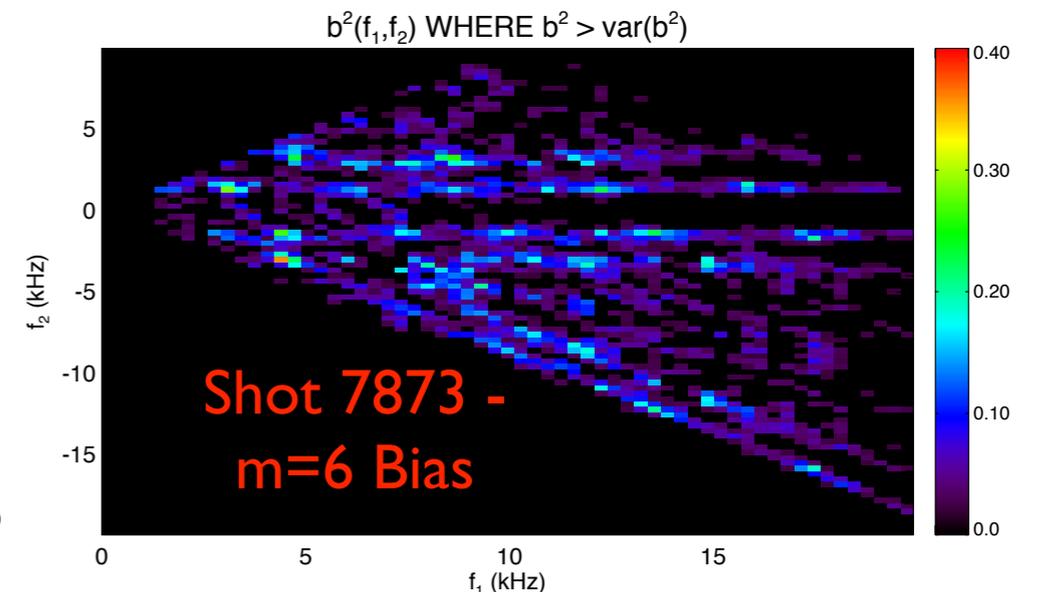
$$\hat{B}(\omega_1, \omega_2) = \langle \hat{S}(\omega_1) \hat{S}(\omega_2) \hat{S}^*(\omega_1 + \omega_2) \rangle$$

$$\hat{b}^2(\omega_1, \omega_2) = \frac{|\hat{B}(\omega_1, \omega_2)|^2}{|\langle \hat{S}(\omega_1) \hat{S}(\omega_2) \rangle|^2 |\langle \hat{S}(\omega_1 + \omega_2) \rangle|^2}$$

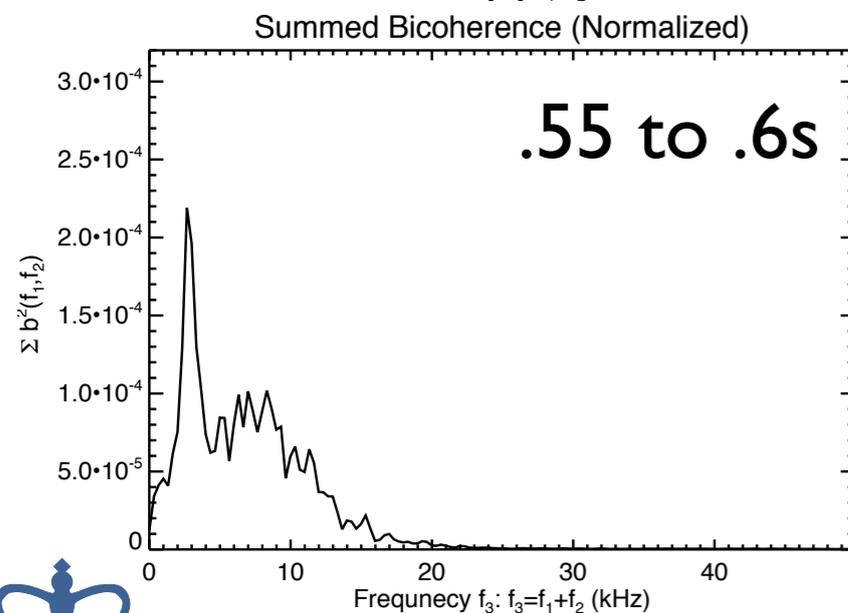
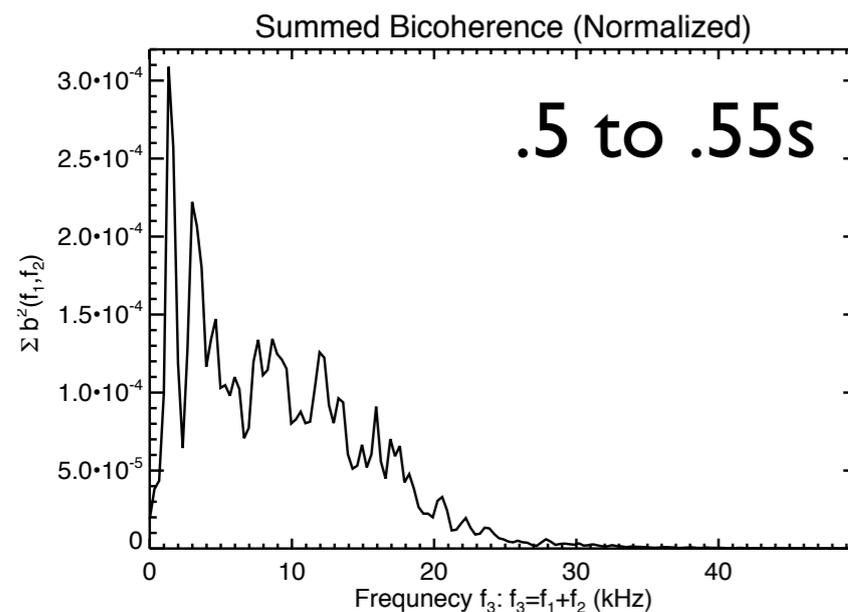
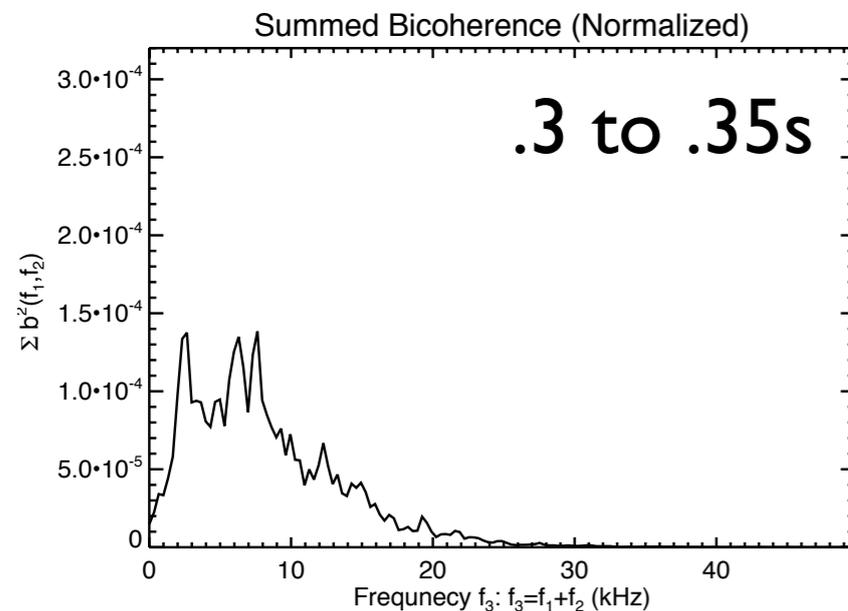


# High Levels of Nonlinear Coupling in Nonaxisymmetric Bias

- Highest level of quadratic coupling see in  $m=6$  bias shot
- See the coupling is very broadband across the spectrum
- Axisymmetric bias displays much lower coupling



# Nonlinear Coupling Increases During Bias



- Summed bicoherence shows the total amount of coupling at a specific frequency
- Shot 7873, m=6 bias
- See a large peak develop through bias and then relax to lower amplitude and higher frequency after bias removed
- Not seen in during axisymmetric bias or unbiased cases

# Conclusions

- Axisymmetric biasing drives the centrifugal interchange mode, even in a turbulent plasma.
- Electrostatic biasing increases coherence length and decreases broadband turbulent fluctuations.
- When nonaxisymmetric bias is applied, the evolution of the turbulent spectrum directly demonstrates the inverse energy cascade. This is the first demonstration of active bi-dimensional turbulence drive in a magnetized plasma.



# Thank You



# Clear identification of mode numbers during bias

- Averaging the phases during the peaks in the coherence identified we can easily trace mode numbers.
- High density phase is less clear (corresponding to lower coherence), but see  $m=3$  and  $m=4$ . The phase evolution of at these frequencies produces clear modes, unlike lower frequencies that are presumably  $m=1$  and  $m=2$  modes.
- In bias we can pick out modes  $m=1,2,3,4,5$  with maybe a hint of 6

